

5. A. HOW FAR FROM THE CENTER OF THE EARTH DOES A SATELLITE TO BE TO HAVE AN ORBITAL PERIOD OF 3 HOURS?

5. A. HOW FAR FROM THE CENTER OF THE EARTH DOES A SATELLITE TO BE TO HAVE AN ORBITAL PERIOD OF 3 HOURS?

UNDERSTANDING THE ORBITAL MECHANICS OF SATELLITES IS FUNDAMENTAL IN FIELDS RANGING FROM TELECOMMUNICATIONS TO SPACE EXPLORATION. ONE INTRIGUING QUESTION THAT OFTEN ARISES IS: HOW FAR FROM THE CENTER OF THE EARTH MUST A SATELLITE BE TO HAVE AN ORBITAL PERIOD OF EXACTLY 3 HOURS? THIS ARTICLE DELVES INTO THE PHYSICS BEHIND SATELLITE ORBITS, PROVIDING A COMPREHENSIVE EXPLANATION AND STEP-BY-STEP CALCULATIONS TO ANSWER THIS QUESTION ACCURATELY.

INTRODUCTION TO SATELLITE ORBITS AND ORBITAL PERIODS

BEFORE EXPLORING THE SPECIFIC CASE OF A SATELLITE WITH A 3-HOUR ORBITAL PERIOD, IT'S ESSENTIAL TO GRASP THE BASIC CONCEPTS OF ORBITAL MECHANICS.

WHAT IS AN ORBITAL PERIOD?

THE ORBITAL PERIOD OF A SATELLITE IS THE TIME IT TAKES TO COMPLETE ONE FULL REVOLUTION AROUND THE EARTH. IT IS TYPICALLY MEASURED IN HOURS, MINUTES, OR SECONDS. THE PERIOD DEPENDS PRIMARILY ON THE SATELLITE'S ALTITUDE (DISTANCE FROM EARTH'S CENTER), THE MASS OF THE EARTH, AND THE GRAVITATIONAL FORCES AT PLAY.

KEPLER'S LAWS AND ORBITAL MECHANICS

KEPLER'S THIRD LAW RELATES THE ORBITAL PERIOD OF A SATELLITE TO ITS ORBITAL RADIUS. SIMPLIFIED FOR EARTH SATELLITES, THE LAW STATES:

$$T^2 = \frac{4\pi^2 R^3}{GM}$$

WHERE:

- T = ORBITAL PERIOD (SECONDS),
- R = DISTANCE FROM EARTH'S CENTER TO THE SATELLITE (METERS),
- G = GRAVITATIONAL CONSTANT ($6.67430 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$),
- M = MASS OF THE EARTH ($5.972 \times 10^{24} \text{ kg}$).

THIS FUNDAMENTAL RELATION ALLOWS US TO DETERMINE THE ORBITAL RADIUS FOR ANY DESIRED PERIOD.

CALCULATING THE REQUIRED ORBITAL RADIUS FOR A 3-HOUR PERIOD

TO FIND HOW FAR THE SATELLITE MUST BE FROM EARTH'S CENTER, WE NEED TO CONVERT THE PERIOD INTO SECONDS AND THEN APPLY KEPLER'S THIRD LAW.

STEP 1: CONVERT PERIOD TO SECONDS

SINCE 1 HOUR = 3600 SECONDS, A 3-HOUR PERIOD IS:

$$T = 3 \times 3600 = 10,800 \text{ SECONDS}$$

STEP 2: REARRANGE KEPLER'S THIRD LAW

REARRANGED TO SOLVE FOR (R) :

$$[R^3 = \frac{GM T^2}{4 \pi^2}]$$

AND THUS:

$$[R = \left(\frac{GM T^2}{4 \pi^2} \right)^{1/3}]$$

STEP 3: PLUG IN KNOWN VALUES

LET'S SUBSTITUTE THE KNOWN CONSTANTS:

- $(G = 6.67430 \times 10^{-11} \text{, } \text{M}^3 \text{kg}^{-1} \text{s}^{-2})$,
- $(M = 5.972 \times 10^{24} \text{, } \text{MATHRM{KG}})$,
- $(T = 10,800 \text{, } \text{MATHRM{S}})$.

CALCULATING NUMERATOR:

$$[GM T^2 = (6.67430 \times 10^{-11}) \times (5.972 \times 10^{24}) \times (10,800)^2]$$

FIRST, COMPUTE (GM) :

$$[GM = 6.67430 \times 10^{-11} \times 5.972 \times 10^{24} \approx 3.986 \times 10^{14} \text{, } \text{MATHRM{M}^3/\text{s}^2}]$$

NEXT, (T^2) :

$$[(10,800)^2 = 116,640,000 \text{, } \text{TEXT{S}}^2]$$

NOW, NUMERATOR:

$$[3.986 \times 10^{14} \times 1.1664 \times 10^8 \approx 4.648 \times 10^{22}]$$

CALCULATE DENOMINATOR:

$$[4 \pi^2 \approx 39.478]$$

FINALLY, COMPUTE (R^3) :

$$[R^3 = \frac{4.648 \times 10^{22}}{39.478} \approx 1.177 \times 10^{21}]$$

NOW, TAKE THE CUBE ROOT:

$$[R = \left(1.177 \times 10^{21} \right)^{1/3}]$$

USING LOGARITHMS OR A CALCULATOR:

$$[R \approx 1.07 \times 10^7 \text{, } \text{TEXT{METERS}}]$$

THIS IS APPROXIMATELY:

$$[R \approx 10,700 \text{, } \text{TEXT{KM}}]$$

INTERPRETING THE RESULT: SATELLITE'S DISTANCE FROM EARTH'S CENTER

THE CALCULATED RADIUS OF APPROXIMATELY 10,700 KM INDICATES THE DISTANCE FROM EARTH'S CENTER WHERE A SATELLITE MUST ORBIT TO HAVE A 3-HOUR PERIOD.

COMPARISON WITH EARTH'S RADIUS

EARTH'S MEAN RADIUS IS ROUGHLY:

$$[R_E \approx 6,371 \text{ km}]$$

THEREFORE, THE ALTITUDE ABOVE EARTH'S SURFACE (h) IS:

$$[h = R - R_E \approx 10,700 \text{ km} - 6,371 \text{ km} \approx 4,329 \text{ km}]$$

THIS MEANS THE SATELLITE WOULD BE APPROXIMATELY 4,329 KM ABOVE EARTH'S SURFACE.

IMPLICATIONS AND TYPES OF SATELLITE ORBITS

UNDERSTANDING WHERE THIS SATELLITE WOULD BE LOCATED HELPS IN CATEGORIZING ITS ORBIT AND UNDERSTANDING ITS MISSION PROFILE.

TYPES OF ORBITS BASED ON ALTITUDE

- LOW EARTH ORBIT (LEO): TYPICALLY BETWEEN 160 KM AND 2,000 KM ALTITUDE.
- MEDIUM EARTH ORBIT (MEO): RANGES FROM ABOUT 2,000 KM TO 35,786 KM.
- GEOSTATIONARY ORBIT (GEO): APPROXIMATELY 35,786 KM ALTITUDE, WITH A PERIOD OF 24 HOURS.

THE SATELLITE IN QUESTION, AT ROUGHLY 4,329 KM ALTITUDE, RESIDES IN THE MEO REGION.

RELEVANCE FOR SATELLITE APPLICATIONS

- SATELLITES AT THIS ALTITUDE CAN BE USED FOR NAVIGATION (LIKE GPS), COMMUNICATIONS, OR SCIENTIFIC OBSERVATIONS.
- THE ORBITAL PERIOD INFLUENCES COVERAGE, DATA TRANSMISSION WINDOWS, AND LATENCY.

ADDITIONAL FACTORS AFFECTING SATELLITE ORBITS

WHILE THE CALCULATIONS ABOVE PROVIDE A THEORETICAL RADIUS BASED ON IDEAL PHYSICS, REAL-WORLD CONDITIONS CAN INFLUENCE SATELLITE ORBITS.

ATMOSPHERIC DRAG AND PERTURBATIONS

- AT LOWER ALTITUDES, ATMOSPHERIC DRAG CAN CAUSE GRADUAL ORBITAL DECAY.
- AT HIGHER ALTITUDES LIKE 4,329 KM, DRAG IS MINIMAL, BUT GRAVITATIONAL PERTURBATIONS FROM THE MOON, SUN, OR EARTH'S OBLATENESS CAN AFFECT THE ORBIT.

ORBITAL STABILITY AND MISSION PLANNING

- ENGINEERS CONSIDER THESE FACTORS WHEN DESIGNING SATELLITE MISSIONS TO ENSURE STABILITY AND LONGEVITY.

CONCLUSION

IN SUMMARY, FOR A SATELLITE TO HAVE AN ORBITAL PERIOD OF EXACTLY 3 HOURS, IT MUST BE APPROXIMATELY 10,700 KM FROM EARTH'S CENTER, WHICH CORRESPONDS TO AN ALTITUDE OF ABOUT 4,329 KM ABOVE EARTH'S SURFACE. THIS POSITION PLACES THE SATELLITE IN THE MEDIUM EARTH ORBIT ZONE, MAKING IT SUITABLE FOR VARIOUS COMMUNICATION AND NAVIGATION APPLICATIONS. UNDERSTANDING THESE CALCULATIONS IS CRUCIAL FOR SATELLITE MISSION DESIGN, ORBITAL MECHANICS ANALYSIS, AND SPACE MISSION PLANNING.

KEY TAKEAWAYS:

- THE ORBITAL PERIOD IS DIRECTLY RELATED TO THE ORBITAL RADIUS THROUGH KEPLER'S THIRD LAW.
- A 3-HOUR ORBIT CORRESPONDS TO A RADIUS OF ROUGHLY 10,700 KM FROM EARTH'S CENTER.
- THE ALTITUDE FROM EARTH'S SURFACE AT THIS RADIUS IS APPROXIMATELY 4,329 KM.
- SUCH SATELLITES ARE IN MEDIUM EARTH ORBIT, WITH SPECIFIC APPLICATIONS DEPENDING ON THEIR ALTITUDE AND PERIOD.

BY MASTERING THESE CONCEPTS, SCIENTISTS AND ENGINEERS CAN ACCURATELY DETERMINE SATELLITE PARAMETERS ESSENTIAL FOR SUCCESSFUL SPACE OPERATIONS AND ADVANCEMENTS IN SATELLITE TECHNOLOGY.

FREQUENTLY ASKED QUESTIONS

HOW DO YOU CALCULATE THE DISTANCE OF A SATELLITE FROM THE EARTH'S CENTER FOR A SPECIFIC ORBITAL PERIOD?

YOU USE KEPLER'S THIRD LAW, WHICH RELATES THE ORBITAL PERIOD TO THE RADIUS OF THE ORBIT. THE FORMULA INVOLVES THE EARTH'S GRAVITATIONAL PARAMETER AND THE DESIRED PERIOD TO FIND THE ORBITAL RADIUS.

WHAT IS THE GRAVITATIONAL PARAMETER OF EARTH USED IN ORBITAL CALCULATIONS?

THE EARTH'S STANDARD GRAVITATIONAL PARAMETER (μ) IS APPROXIMATELY $3.986 \times 10^{14} \text{ m}^3/\text{s}^2$.

HOW IS THE ORBITAL RADIUS RELATED TO THE SATELLITE'S ORBITAL PERIOD?

THE ORBITAL RADIUS (DISTANCE FROM EARTH'S CENTER) IS PROPORTIONAL TO THE PERIOD RAISED TO THE POWER OF TWO-THIRDS, BASED ON KEPLER'S THIRD LAW.

WHAT IS THE FORMULA TO FIND THE ORBITAL RADIUS FOR A SATELLITE WITH A 3-HOUR PERIOD?

$R = \left[\frac{\mu T^2}{4\pi^2} \right]^{1/3}$, WHERE T IS THE PERIOD IN SECONDS.

CONVERT 3 HOURS INTO SECONDS FOR ORBITAL CALCULATIONS.

3 HOURS EQUALS 10,800 SECONDS.

WHAT IS THE APPROXIMATE DISTANCE FROM EARTH'S CENTER FOR A SATELLITE WITH A

3-HOUR PERIOD?

USING THE FORMULA, THE SATELLITE WOULD BE APPROXIMATELY 42,000 KILOMETERS FROM EARTH'S CENTER.

IS A SATELLITE WITH A 3-HOUR ORBITAL PERIOD CONSIDERED LOW EARTH ORBIT OR HIGHER ORBIT?

AN ORBIT WITH A 3-HOUR PERIOD IS HIGHER THAN TYPICAL LOW EARTH ORBIT; IT'S CLOSER TO MEDIUM EARTH ORBIT ALTITUDES.

WHY IS KNOWING THE ORBITAL RADIUS IMPORTANT FOR SATELLITE DEPLOYMENT?

KNOWING THE ORBITAL RADIUS HELPS IN DETERMINING SATELLITE COVERAGE, COMMUNICATION RANGE, AND MISSION PLANNING FOR DESIRED ORBITAL CHARACTERISTICS.

HOW DOES THE ORBITAL PERIOD AFFECT SATELLITE SPEED AND ALTITUDE?

A LONGER ORBITAL PERIOD CORRESPONDS TO A HIGHER ALTITUDE AND LOWER ORBITAL SPEED, WHILE A SHORTER PERIOD INDICATES A LOWER ALTITUDE AND HIGHER SPEED.

ADDITIONAL RESOURCES

HOW FAR FROM THE CENTER OF THE EARTH DOES A SATELLITE TO BE TO HAVE AN ORBITAL PERIOD OF 3 HOURS?

UNDERSTANDING THE PRECISE ALTITUDE AT WHICH A SATELLITE MUST ORBIT TO COMPLETE A FULL REVOLUTION AROUND THE EARTH IN EXACTLY THREE HOURS IS A FASCINATING APPLICATION OF ORBITAL MECHANICS. WHEN WE ASK, "HOW FAR FROM THE CENTER OF THE EARTH DOES A SATELLITE HAVE TO BE TO HAVE AN ORBITAL PERIOD OF 3 HOURS?", WE ARE ESSENTIALLY SEEKING THE ORBITAL RADIUS THAT SATISFIES THE GRAVITATIONAL AND KINEMATIC CONDITIONS FOR THIS SPECIFIC PERIOD. THIS QUESTION INVOLVES FUNDAMENTAL CONCEPTS SUCH AS NEWTON'S LAW OF GRAVITATION, CIRCULAR ORBITAL MOTION, AND THE RELATIONSHIP BETWEEN ORBITAL RADIUS AND PERIOD.

IN THIS ARTICLE, WE WILL EXPLORE THE DETAILED PROCESS OF CALCULATING THIS DISTANCE, STARTING FROM THE BASIC PRINCIPLES, WORKING THROUGH THE RELEVANT EQUATIONS, AND DISCUSSING THE IMPLICATIONS OF THE RESULTS. WHETHER YOU'RE AN AEROSPACE ENTHUSIAST, A STUDENT, OR A PROFESSIONAL IN THE FIELD, THIS COMPREHENSIVE GUIDE WILL DEEPEN YOUR UNDERSTANDING OF HOW ORBITAL PERIODS RELATE TO SATELLITE ALTITUDES.

THE FUNDAMENTALS OF ORBITAL MOTION

BEFORE DIVING INTO CALCULATIONS, IT'S ESSENTIAL TO UNDERSTAND THE CORE PHYSICS PRINCIPLES GOVERNING SATELLITE ORBITS.

NEWTON'S LAW OF UNIVERSAL GRAVITATION

AT THE HEART OF ORBITAL MECHANICS IS NEWTON'S LAW OF UNIVERSAL GRAVITATION, WHICH STATES THAT THE GRAVITATIONAL FORCE (F) BETWEEN TWO MASSES (M) (EARTH) AND (m) (SATELLITE) SEPARATED BY A DISTANCE (r) IS:

$$F = G \frac{Mm}{r^2}$$

WHERE:

- (G) IS THE GRAVITATIONAL CONSTANT, APPROXIMATELY $(6.674 \times 10^{-11} \text{ Nm}^2/\text{kg}^2)$,

- (M) IS THE MASS OF EARTH, APPROXIMATELY $(5.972 \times 10^{24} \text{ kg})$,
- (m) IS THE MASS OF THE SATELLITE,
- (r) IS THE DISTANCE FROM EARTH'S CENTER TO THE SATELLITE.

CIRCULAR ORBITAL VELOCITY AND PERIOD

FOR A SATELLITE IN A CIRCULAR ORBIT, THE GRAVITATIONAL FORCE PROVIDES THE NECESSARY CENTRIPETAL FORCE:

$$G \frac{M m}{r^2} = \frac{m v^2}{r}$$

WHICH SIMPLIFIES TO:

$$v = \sqrt{\frac{G M}{r}}$$

WHERE (v) IS THE ORBITAL VELOCITY.

THE ORBITAL PERIOD (T) , THE TIME TAKEN TO COMPLETE ONE FULL ORBIT, RELATES TO THE ORBITAL RADIUS AND VELOCITY:

$$T = \frac{2 \pi r}{v}$$

SUBSTITUTING (v) :

$$T = 2 \pi r \sqrt{\frac{r}{G M}} = 2 \pi \sqrt{\frac{r^3}{G M}}$$

THIS IS A KEY RELATIONSHIP CONNECTING ORBITAL PERIOD AND RADIUS.

DERIVING THE ORBITAL RADIUS FOR A 3-HOUR PERIOD

GIVEN THE ORBITAL PERIOD (T) IS 3 HOURS, OR IN SECONDS:

$$T = 3 \times 3600 = 10,800 \text{ seconds}$$

OUR GOAL IS TO FIND THE ORBITAL RADIUS (r) .

REARRANGED ORBITAL PERIOD EQUATION

STARTING FROM:

$$T = 2 \pi \sqrt{\frac{r^3}{G M}}$$

SQUARING BOTH SIDES:

$$T^2 = 4 \pi^2 \frac{r^3}{G M}$$

SOLVING FOR (R^3) :

$$R^3 = \frac{G M T^2}{4 \pi^2}$$

AND THEN TAKING THE CUBE ROOT:

$$R = \left(\frac{G M T^2}{4 \pi^2} \right)^{1/3}$$

PLUGGING IN THE VALUES

USING THE KNOWN CONSTANTS:

- $(G = 6.674 \times 10^{-11} \text{ , } \text{Nm}^2/\text{kg}^2 \text{)}$,
- $(M = 5.972 \times 10^{24} \text{ , } \text{kg} \text{)}$,
- $(T = 10,800 \text{ , } \text{s} \text{)}$.

CALCULATE NUMERATOR:

$$G M T^2 = (6.674 \times 10^{-11}) \times (5.972 \times 10^{24}) \times (10,800)^2$$

FIRST, MULTIPLY (G) AND (M) :

$$6.674 \times 10^{-11} \times 5.972 \times 10^{24} \approx 3.986 \times 10^{14}$$

NEXT, COMPUTE (T^2) :

$$(10,800)^2 = 116,640,000$$

NOW, MULTIPLY:

$$3.986 \times 10^{14} \times 1.1664 \times 10^8 \approx 4.651 \times 10^{22}$$

CALCULATE DENOMINATOR:

$$4 \pi^2 \approx 39.478$$

FINALLY, COMPUTE (R^3) :

$$R^3 = \frac{4.651 \times 10^{22}}{39.478} \approx 1.177 \times 10^{21}$$

NOW, TAKE THE CUBE ROOT:

$$R =$$

$$R = \left(1.177 \times 10^{21} \right)^{1/3}$$

RECALL THAT:

$$\left(10^{21} \right)^{1/3} = 10^7$$

AND:

$$(1.177)^{1/3} \approx 1.055$$

So,

$$R \approx 1.055 \times 10^7 \text{ meters}$$

WHICH IS APPROXIMATELY:

$$R \approx 10,550,000 \text{ meters}$$

OR ABOUT 10,550 KM FROM THE EARTH'S CENTER.

INTERPRETING THE RESULTS

THIS CALCULATION INDICATES THAT A SATELLITE MUST ORBIT APPROXIMATELY 10,550 KM FROM EARTH'S CENTER TO HAVE AN ORBITAL PERIOD OF EXACTLY THREE HOURS.

COMPARING TO EARTH'S RADIUS

EARTH'S MEAN RADIUS (R_E) IS APPROXIMATELY 6,371 KM. THEREFORE, THE SATELLITE'S ALTITUDE (h) ABOVE EARTH'S SURFACE IS:

$$h = R - R_E = 10,550 \text{ km} - 6,371 \text{ km} \approx 4,179 \text{ km}$$

THIS MEANS THE SATELLITE WOULD ORBIT ROUGHLY 4,179 KM ABOVE EARTH'S SURFACE.

PRACTICAL IMPLICATIONS AND CONTEXT

SATELLITE ORBITS AND PERIODS

- LOW EARTH ORBIT (LEO): TYPICALLY RANGES FROM ABOUT 160 KM TO 2,000 KM ALTITUDE, WITH ORBITAL PERIODS AROUND 90 TO 120 MINUTES.
- MEDIUM EARTH ORBIT (MEO): RANGES FROM 2,000 KM TO ABOUT 35,786 KM, WITH PERIODS FROM ABOUT 2 TO 24 HOURS.
- GEOSTATIONARY ORBIT: AT APPROXIMATELY 35,786 KM ALTITUDE, WITH A PERIOD OF 24 HOURS.

A 3-HOUR ORBIT PLACES THE SATELLITE IN THE MEO REGION, CLOSER TO THE LOWER END OF THAT SPECTRUM, BUT STILL SIGNIFICANTLY ABOVE LEO.

RELEVANCE TO SATELLITE MISSIONS

SATELLITES WITH PERIODS OF AROUND 3 HOURS ARE OFTEN USED FOR:

- NAVIGATION AND POSITIONING SYSTEMS.
- COMMUNICATION RELAYS.
- SCIENTIFIC OBSERVATIONS REQUIRING INTERMEDIATE ALTITUDES.

LIMITATIONS AND CONSIDERATIONS

- ORBITAL DECAY: AT SUCH ALTITUDES, ATMOSPHERIC DRAG IS MINIMAL BUT NOT ZERO, ESPECIALLY AT LOWER PERIGEEES.
- RADIATION EXPOSURE: HIGHER ALTITUDES EXPOSE SATELLITES TO MORE RADIATION, IMPACTING DESIGN CHOICES.
- LAUNCH REQUIREMENTS: REACHING THIS ORBIT DEMANDS SPECIFIC LAUNCH VEHICLES AND ENERGY CONSIDERATIONS.

ADDITIONAL FACTORS IN REAL-WORLD SCENARIOS

WHILE THE CALCULATIONS ABOVE ASSUME A PERFECT CIRCULAR ORBIT AND NEGLECT PERTURBATIONS, REAL SATELLITES EXPERIENCE VARIOUS FORCES AND INFLUENCES.

EARTH'S OBLATENESS AND GRAVITY ANOMALIES

- EARTH'S EQUATORIAL BULGE CAUSES SLIGHT VARIATIONS IN GRAVITATIONAL PULL, AFFECTING ORBITAL STABILITY.
- MANY SATELLITES REQUIRE STATION-KEEPING MANEUVERS TO MAINTAIN THEIR ORBITS.

ATMOSPHERIC DRAG AND SOLAR RADIATION PRESSURE

- EVEN AT SEVERAL THOUSAND KILOMETERS ALTITUDE, RESIDUAL ATMOSPHERIC PARTICLES AND SOLAR RADIATION CAN INFLUENCE ORBIT OVER TIME.

ORBITAL INCLINATION AND ECCENTRICITY

- ACTUAL SATELLITE ORBITS ARE OFTEN INCLINED AND ELLIPTICAL, WHICH AFFECTS PERIOD AND ALTITUDE RELATIONSHIPS.

SUMMARY AND FINAL THOUGHTS

IN CONCLUSION, A SATELLITE MUST BE APPROXIMATELY 10,550 KM FROM EARTH'S CENTER TO HAVE AN ORBITAL PERIOD OF 3 HOURS. THIS CALCULATION RELIES ON FUNDAMENTAL PHYSICS PRINCIPLES AND PROVIDES A FOUNDATIONAL UNDERSTANDING OF HOW ORBITAL PERIOD RELATES TO DISTANCE.

KEY TAKEAWAYS:

- THE ORBITAL RADIUS FOR A 3-HOUR PERIOD IS ABOUT 10,550 KM FROM EARTH'S CENTER.
- THE CORRESPONDING ALTITUDE ABOVE EARTH'S SURFACE IS APPROXIMATELY 4,179 KM.
- THESE CALCULATIONS ARE IDEALIZED; REAL-WORLD CONDITIONS INTRODUCE COMPLEXITIES.

UNDERSTANDING THESE RELATIONSHIPS IS CRUCIAL FOR DESIGNING SATELLITE MISSIONS, PREDICTING ORBITAL BEHAVIORS, AND ANALYZING SATELLITE COVERAGE AND CAPABILITIES. AS TECHNOLOGY ADVANCES, PRECISE ORBITAL CALCULATIONS LIKE THIS UNDERPIN THE SUCCESSFUL DEPLOYMENT AND OPERATION OF COUNTLESS SPACE-BASED SYSTEMS THAT MODERN LIFE RELIES UPON.

REFERENCES AND FURTHER READING

- "ORBITAL MECHANICS FOR ENGINEERING STUDENTS" BY HOWARD D. CURTIS
- NASA'S FUNDAMENTALS OF ASTRODYNAMICS
- "INTRODUCTION TO SPACE DYNAMICS" BY WILLIAM TYRRELL THOMSON
- ONLINE CALCULATORS FOR ORBITAL PERIOD AND RADIUS

FEEL FREE TO EXPLORE THESE RESOURCES FOR A

5 A How Far From The Center Of The Earth Does A Satellite To Be To Have An Orbital Period Of 3 Hours

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