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Introduction to Compton Scattering

Compton scattering is a fundamental phenomenon in quantum physics and particle interactions, describing the inelastic scattering of a photon by a charged particle, typically an electron. When a high-energy photon interacts with an electron, it transfers part of its energy to the electron, resulting in a decrease in the photon's energy and a corresponding increase in the electron's kinetic energy. This process provides essential insights into the particle nature of light and the structure of matter.

In this particular problem, a photon with an initial energy of 0.1 MeV (mega-electron volts) undergoes Compton scattering at an angle of 60 degrees. The goal is to determine the kinetic energy gained by the electron post-collision. To achieve this, we need to understand the fundamental principles governing Compton scattering, the relevant formulas, and step-by-step calculations.

Understanding the Initial Conditions

Photon Energy and Wavelength

The initial photon energy is given as:

$$E_i = 0.1 \text{ MeV}$$

Since energy and wavelength are related via the photon's energy:

$$E = \frac{hc}{\lambda}$$

where:

- h is Planck's constant ($6.626 \times 10^{-34} \text{ Js}$),
- c is the speed of light ($3.0 \times 10^8 \text{ m/s}$),
- λ is the wavelength.

Converting the initial energy from MeV to Joules:

$$1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$$

thus,

$$0.1 \text{ MeV} = 0.1 \times 10^6 \text{ eV} = 1.602 \times 10^{-14} \text{ J}$$

The initial photon wavelength:

$$\lambda_i = \frac{hc}{E_i} = \frac{6.626 \times 10^{-34} \times 3.0 \times 10^8}{1.602 \times 10^{-14}} \approx 1.24 \times 10^{-12} \text{ m}$$

Compton Scattering Fundamentals

Compton Wavelength Shift Formula

The core formula for the change in photon wavelength due to scattering is:

$$\Delta \lambda = \lambda_f - \lambda_i = \frac{h}{m_e c} (1 - \cos \theta)$$

where:

- λ_f is the final wavelength,
- m_e is the rest mass of the electron ($9.109 \times 10^{-31} \text{ kg}$),
- θ is the scattering angle of the photon.

The term $\frac{h}{m_e c}$ is known as the Compton wavelength of the electron, approximately:

$$\lambda_C = \frac{h}{m_e c} \approx 2.43 \times 10^{-12} \text{ m}$$

Given the scattering angle $\theta = 60^\circ$:

$$\cos 60^\circ = 0.5$$

Thus, the wavelength shift:

$$\Delta \lambda = 2.43 \times 10^{-12} \times (1 - 0.5) = 2.43 \times 10^{-12} \times 0.5 = 1.215 \times 10^{-12} \text{ m}$$

Calculating the Final Photon Energy and Wavelength

The final photon wavelength after scattering:

$$\lambda_f = \lambda_i + \Delta \lambda = 1.24 \times 10^{-12} + 1.215 \times 10^{-12} = 2.455 \times 10^{-12} \text{ m}$$

Corresponding final photon energy:

$$E_f = \frac{hc}{\lambda_f} = \frac{6.626 \times 10^{-34} \times 3.0 \times 10^8}{2.455 \times 10^{-12}} \approx 8.11 \times 10^{-14} \text{ J}$$

\]

Converting back to eV:

\[

$$E_f \approx \frac{8.11 \times 10^{-14}}{1.602 \times 10^{-19}} \approx 0.507, \text{MeV}$$

\]

Note: Since the initial energy was 0.1 MeV, this calculation indicates an inconsistency — the energy should decrease after scattering, not increase. The mistake arises from interpreting the wavelength shift directly; instead, we should use the correct energy relation considering the inverse proportionality between energy and wavelength.

Correct Approach: Using Compton Formula for Photons

The more precise way to determine the energy after scattering involves the Compton formula expressed in terms of photon energies:

\[

$$E_f = \frac{E_i}{1 + \frac{E_i}{m_e c^2}(1 - \cos \theta)}$$

\]

where:

- $(m_e c^2)$ is the rest energy of the electron ($(0.511, \text{MeV})$),
- (E_i) is the initial photon energy,
- (θ) is the scattering angle.

Given:

\[

$$E_i = 0.1, \text{MeV}$$

\]

\[

$$\theta = 60^\circ$$

\]

\[

$$\cos 60^\circ = 0.5$$

\]

Calculating the denominator:

$$D = 1 + \frac{0.1}{0.511} \times (1 - 0.5) = 1 + 0.1957 \times 0.5 = 1 + 0.0979 = 1.0979$$

Final photon energy:

$$E_f = \frac{0.1}{1.0979} \approx 0.091, \text{ MeV}$$

Determining the Electron's Kinetic Energy

The energy transferred to the electron (i.e., the kinetic energy (K)) is:

$$K = E_i - E_f$$

Substituting the calculated values:

$$K = 0.1, \text{ MeV} - 0.091, \text{ MeV} = 0.009, \text{ MeV}$$

Expressed in eV:

$$K = 0.009 \times 10^6 = 9000, \text{ eV}$$

Thus, the electron gains approximately 9 keV of kinetic energy after the scattering.

Summary of Results

- Initial photon energy: 0.1 MeV
- Scattering angle: 60°
- Final photon energy: approximately 0.091 MeV
- Electron's kinetic energy: approximately 0.009 MeV or 9 keV

Implications and Significance

This calculation exemplifies how photon-electron interactions depend critically on scattering angles and initial photon energies. The transfer of energy in Compton scattering directly correlates with the scattering angle, with larger angles resulting in more significant energy transfer to electrons.

The obtained kinetic energy (around 9 keV) indicates that high-energy photons like 0.1 MeV can impart substantial energy to electrons, which is instrumental in numerous applications such as:

- Medical imaging (e.g., X-ray imaging)
- Radiation therapy
- Particle physics experiments
- Understanding cosmic phenomena involving high-energy photons

Additional Considerations

- Limitations: The calculations assume free electrons at rest and neglect other potential interactions.
- Relativistic Effects: At energies around 0.1 MeV, relativistic effects are significant, and the formulas used incorporate relativistic physics.
- Experimental Verification: Such theoretical calculations are validated through experimental measurements, confirming the accuracy of quantum electrodynamics predictions.

Conclusion

In conclusion, a photon with an initial energy of 0.1 MeV undergoing Compton scattering at a 60-degree angle transfers approximately 9 keV of kinetic energy to the electron. This process underscores the dual wave-particle nature of light and the importance of quantum mechanics in explaining high-energy photon interactions. Understanding these principles is vital in fields spanning astrophysics, medical physics, and fundamental particle research.

Keywords: Compton scattering, photon energy, electron kinetic energy, scattering angle, relativistic physics, quantum mechanics, photon-electron interaction, energy transfer

Frequently Asked Questions

What is Compton scattering and how does it affect a photon with initial energy of 0.1 MeV at a 60° scattering angle?

Compton scattering is the inelastic scattering of a photon by a charged particle, typically an electron, resulting in a decrease in the photon's energy and a change in its direction. For a photon with initial energy of 0.1 MeV scattered at 60°, the photon loses some energy, which is transferred as kinetic energy to the electron.

How do you calculate the change in wavelength of a photon during Compton scattering at a 60° angle?

The change in wavelength $\Delta\lambda$ is given by the Compton formula: $\Delta\lambda = (h / (m_e c)) (1 - \cos \theta)$. For $\theta = 60^\circ$, $\Delta\lambda \approx 0.0243$ nm, using $h/(m_e c) \approx 0.0243$ nm.

What is the formula to determine the energy of the photon after scattering at a 60° angle?

The final photon energy E' can be found using: $1/E' - 1/E = (1/m_e c^2) (1 - \cos \theta)$, where E is the initial energy. Rearranged, $E' = E / [1 + (E / (m_e c^2)) (1 - \cos \theta)]$.

How do you compute the kinetic energy gained by the electron during

Compton scattering at 60°?

The electron's kinetic energy (K) equals the initial photon energy minus the scattered photon energy: $K = E - E'$. Calculate E' using the Compton formula, then subtract from initial E to find K .

What is the rest energy of the electron, and why is it important in calculating the scattered photon energy?

The rest energy of the electron is approximately 0.511 MeV. It is important because it appears in the Compton formula and helps relate the change in photon energy to the scattering angle, especially at high photon energies like 0.1 MeV.

Can you provide a step-by-step calculation to find the kinetic energy of the electron after scattering?

Yes. First, calculate the scattered photon energy E' using: $E' = E / [1 + (E / 0.511) (1 - \cos 60^\circ)]$. Then, find $K = E - E'$. For $E = 0.1$ MeV, plug in the numbers to get the electron's kinetic energy.

What are the approximate values of the scattered photon energy and the electron's kinetic energy after the scattering?

Using the formulas, the scattered photon energy E' is approximately 0.088 MeV, and the electron's kinetic energy K is approximately 0.012 MeV (12 keV).

Why is understanding Compton scattering important in medical imaging and radiation physics?

Compton scattering explains how high-energy photons interact with matter, affecting image quality and radiation dose calculations in medical imaging techniques like X-ray and gamma-ray diagnostics, as well as in radiation shielding and dosimetry.

Additional Resources

1) A Photon Of Initial Energy 0.1 MeV Undergoes Compton Scattering At An Angle Of 60. Find (a) The Kinetic

Imagine a tiny packet of light, a photon, hurtling through space at near the speed of light. This photon carries a specific amount of energy—0.1 MeV—and, in its journey, it encounters an electron. The interaction between this high-energy photon and the electron results in a phenomenon known as Compton scattering, a fundamental process that reveals the quantum nature of light and matter. In this article, we

will explore the detailed physics behind this process, focusing on calculating the kinetic energy gained by the electron when the photon is scattered at a 60-degree angle.

Understanding Compton Scattering: A Fundamental Quantum Phenomenon

Compton scattering, named after Arthur Compton who first observed it in 1923, is a quantum mechanical process describing the inelastic scattering of a photon by a charged particle, most commonly an electron. Unlike classical scattering, where light is deflected without energy loss, Compton scattering involves an exchange of energy and momentum, leading to an increase in the wavelength (or decrease in energy) of the photon and a corresponding recoil of the electron.

Key Concepts in Compton Scattering:

- Photon Energy and Wavelength: The initial photon has an energy of 0.1 MeV, which corresponds to a specific wavelength via the relation $(E = hc / \lambda)$.
- Inelastic Process: The photon transfers some of its energy to the electron, resulting in the electron gaining kinetic energy.
- Angular Dependence: The amount of energy transferred depends on the scattering angle, which in this case is 60 degrees.

Initial Parameters and Basic Formulas

Before diving into calculations, it is essential to establish the initial parameters of the problem:

- Initial photon energy, (E_i) : 0.1 MeV
- Initial photon wavelength, (λ_i) : Derived from (E_i)
- Scattering angle, (θ) : 60 degrees

Converting Energy to Wavelength:

Using the relation:

$$\lambda_i = \frac{hc}{E_i}$$

\]

where:

- (h) (Planck's constant): (6.626×10^{-34}) Js
- (c) (speed of light): (3.0×10^8) m/s
- (E_i) : $0.1 \text{ MeV} = (0.1 \times 10^6 \times 1.602 \times 10^{-19}) \text{ J} \approx (1.602 \times 10^{-14}) \text{ J}$

Calculating:

\[

$$\lambda_i = \frac{6.626 \times 10^{-34} \times 3.0 \times 10^8}{1.602 \times 10^{-14}} \approx 1.24 \times 10^{-12} \text{ meters}$$

\]

This is a very short wavelength, typical for gamma-ray photons.

The Compton Wavelength Shift Formula:

A central relation in Compton scattering relates the change in photon wavelength to the scattering angle:

\[

$$\Delta \lambda = \lambda_f - \lambda_i = \frac{h}{m_e c} (1 - \cos \theta)$$

\]

where:

- (λ_f) : Final photon wavelength after scattering
- (m_e) : Rest mass of the electron (9.109×10^{-31}) kg
- (θ) : Scattering angle (60 degrees here)

The term $(\frac{h}{m_e c})$ is known as the Compton wavelength of the electron:

\[

$$\lambda_C = \frac{6.626 \times 10^{-34}}{9.109 \times 10^{-31} \times 3.0 \times 10^8} \approx 2.43 \times 10^{-12} \text{ meters}$$

\]

Calculating the Wavelength Shift and Final Photon Energy

Step 1: Determine $\Delta \lambda$

Given $\theta = 60^\circ$:

$$\begin{aligned} & \left[\right. \\ & \cos 60^\circ = 0.5 \\ & \left. \right] \end{aligned}$$

So,

$$\begin{aligned} & \left[\right. \\ & \Delta \lambda = \lambda_C (1 - \cos \theta) = 0.5 \lambda_C = 0.5 \times 2.43 \times 10^{-12} \approx 1.215 \times 10^{-12} \text{ meters} \\ & \left. \right] \end{aligned}$$

Step 2: Calculate Final Wavelength λ_f

$$\begin{aligned} & \left[\right. \\ & \lambda_f = \lambda_i + \Delta \lambda = 1.24 \times 10^{-12} + 1.215 \times 10^{-12} \approx 2.455 \times 10^{-12} \text{ meters} \\ & \left. \right] \end{aligned}$$

Step 3: Find the Final Photon Energy E_f

Using $E = hc / \lambda_f$:

$$\begin{aligned} & \left[\right. \\ & E_f = \frac{6.626 \times 10^{-34} \times 3.0 \times 10^8}{2.455 \times 10^{-12}} \approx 8.11 \times 10^{-14} \text{ Joules} \\ & \left. \right] \end{aligned}$$

Converting back to MeV:

$$\begin{aligned} & \left[\right. \\ & E_f = \frac{8.11 \times 10^{-14}}{1.602 \times 10^{-13}} \approx 0.507 \text{ MeV} \\ & \left. \right] \end{aligned}$$

But note that the initial energy was 0.1 MeV; this suggests an inconsistency because the energy should decrease after scattering. The mistake stems from the over-simplification of the wavelength approach without considering the relativistic energy relation. Let's correct this.

Relativistic Energy and Momentum Relations

The previous step used a direct wavelength approach, which, while illustrative, isn't precise for high-energy photons like 0.1 MeV photons. Instead, we should utilize the energy-momentum relations.

Photon Energy and Momentum:

$$E = pc$$

where:

- p : photon momentum

Expressing the initial photon momentum:

$$p_i = \frac{E_i}{c} = \frac{0.1 \times 10^6 \times 1.602 \times 10^{-19}}{3 \times 10^8} \approx 5.34 \times 10^{-22} \text{ kg}\cdot\text{m/s}$$

Electron Rest Mass Energy:

$$m_e c^2 = 0.511 \text{ MeV}$$

which is significant compared to the photon energy, so relativistic effects are important.

Applying the Compton Scattering Formula for Photons

The more precise relation for the scattered photon energy is:

$$E = \frac{hc}{\lambda}$$

$$E_f = \frac{E_i}{1 + \frac{E_i}{m_e c^2} (1 - \cos \theta)}$$

This formula accounts for relativistic effects and energy conservation.

Calculating E_f :

Given:

- $E_i = 0.1 \text{ MeV}$
- $m_e c^2 = 0.511 \text{ MeV}$
- $\cos 60^\circ = 0.5$

Plugging into the formula:

$$E_f = \frac{0.1}{1 + \frac{0.1}{0.511} (1 - 0.5)} = \frac{0.1}{1 + 0.1957 \times 0.5} = \frac{0.1}{1 + 0.09785} = \frac{0.1}{1.09785} \approx 0.091 \text{ MeV}$$

Thus, the photon's energy after scattering at 60 degrees is approximately 0.091 MeV.

Determining the Electron's Kinetic Energy

The electron gains the difference in energy of the photon:

$$K.E._e = E_i - E_f$$

Substituting values:

$$K.E._e = 0.1 \text{ MeV} - 0.091 \text{ MeV} = 0.009 \text{ MeV}$$

Expressed in more familiar units:

$$K.E._e = 9 \text{ eV}$$

$$0.009 \text{ MeV} = 9 \text{ keV}$$

\]

Final Result:

The electron acquires approximately 9 keV of kinetic energy when the photon is scattered at 60 degrees, starting from an initial energy of 0.1 MeV.

Implications and Significance

This calculation exemplifies how quantum mechanics and relativity intertwine in high-energy physics phenomena. The relatively small transfer of energy—about 9 keV—may seem minor compared to the photon's initial energy, but it has profound implications for understanding interactions at the nuclear and subatomic levels.

Real-world applications include:

- Medical Imaging: Understanding photon-electron interactions helps improve techniques like PET scans and gamma-ray imaging.
- Nuclear Physics: Compton scattering experiments probe the structure of matter.
- Astrophysics: High-energy photon interactions inform models

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